

A High-Order Discontinuous Galerkin Approach for Physics-Based Thermospheric Modeling



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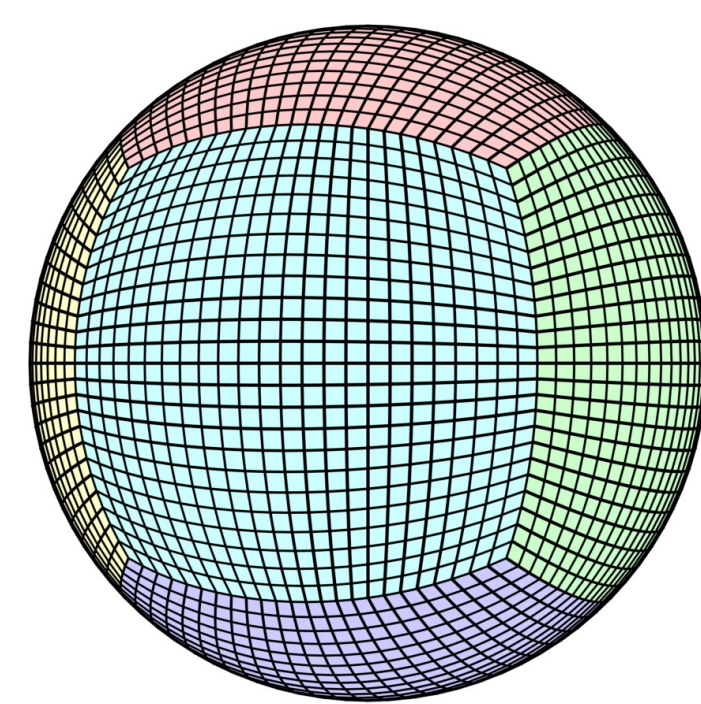
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1. Motivation

- Big interest towards the exploitation of the **Low-Earth Orbiting (LEO)** region of the atmosphere \Rightarrow mega-constellations of satellites.
- Need for accurate **trajectory prediction** in orbits heavily influenced by **atmospheric drag**.
- **Physics-based models** of the ionosphere-thermosphere to improve forecasting capabilities of empirical models.
- Introduce modern **high-order numerical methods** for physics-based thermospheric modeling [1].
 - **High accuracy** at reduced computational cost.
 - Ability to **resolve complex physics**.
 - Multiscale variations, coupled phenomena...

2. Computational Approach

Exasim



- High-order **discontinuous Galerkin (DG)** code [2].
- **Implicit-in-time**: DIRK temporal schemes.
- **Matrix-free** approach, optimized for **GPUs**.
- Automatic code generation.

Ability to **preserve accuracy** in adapted and unstructured meshes.

- **Mesh refinement** to handle complex physics.
- **Cube-sphere meshes** to avoid pole singularities.

3. Physics-Based Model

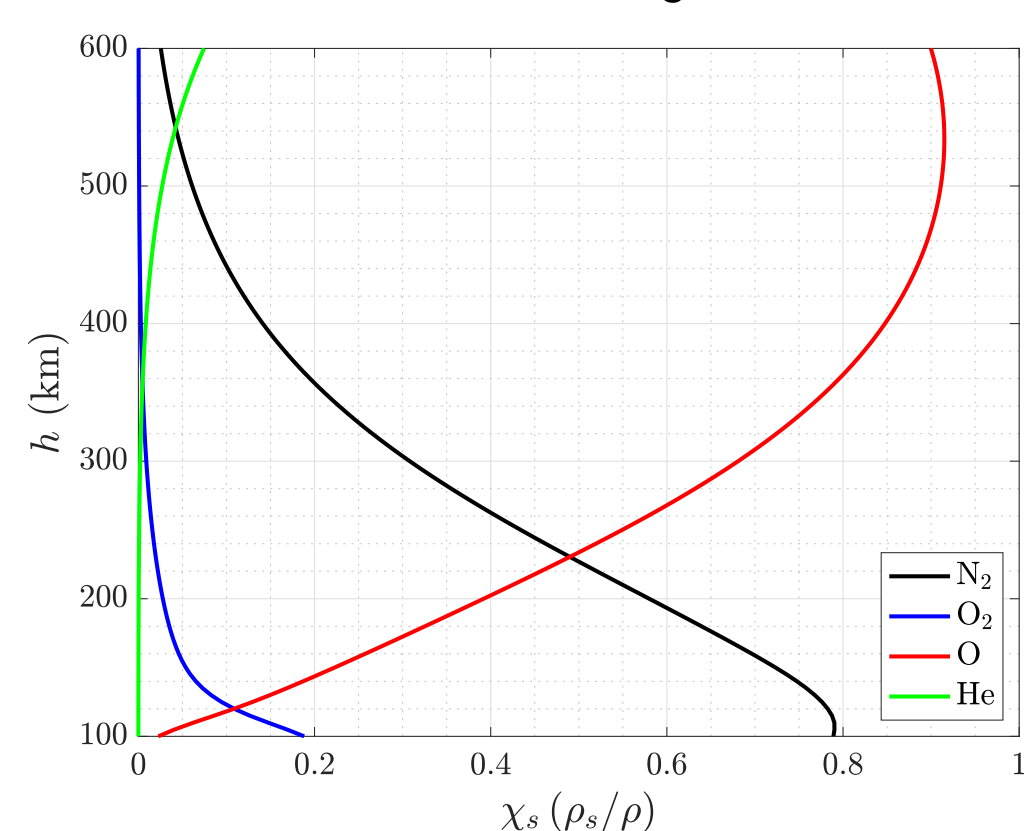
Physics-based thermospheric model consisting of a neutral gas mixture in non-hydrostatic equilibrium.

- Compressible Navier-Stokes equations in the rotating Earth.
- **Non-conservative formulation** to handle multiscale variations.

$$\mathbf{u} = \begin{bmatrix} \log(\rho) \\ \sqrt{\rho} v_i \\ \sqrt{\rho} T \end{bmatrix} \quad \begin{array}{l} \text{➢ Ensure } \textit{positivity} \text{ and deal with low densities.} \\ \text{➢ Scaled variables to } \textit{reduce stiffness} \text{ of the system of equations.} \end{array}$$

- Neutral gas mixture with **chemically frozen composition**.

- A priori knowledge of composition of most relevant species.
- Variable properties with altitude.



- Heating source due to **EUV radiation**.
 - Measures amount of energy being transferred due to **photoionization**.
 - Based on EUVAC model [3], driven by the $F_{10.7}$ index and empirical and measured data of the **EUV spectrum**.

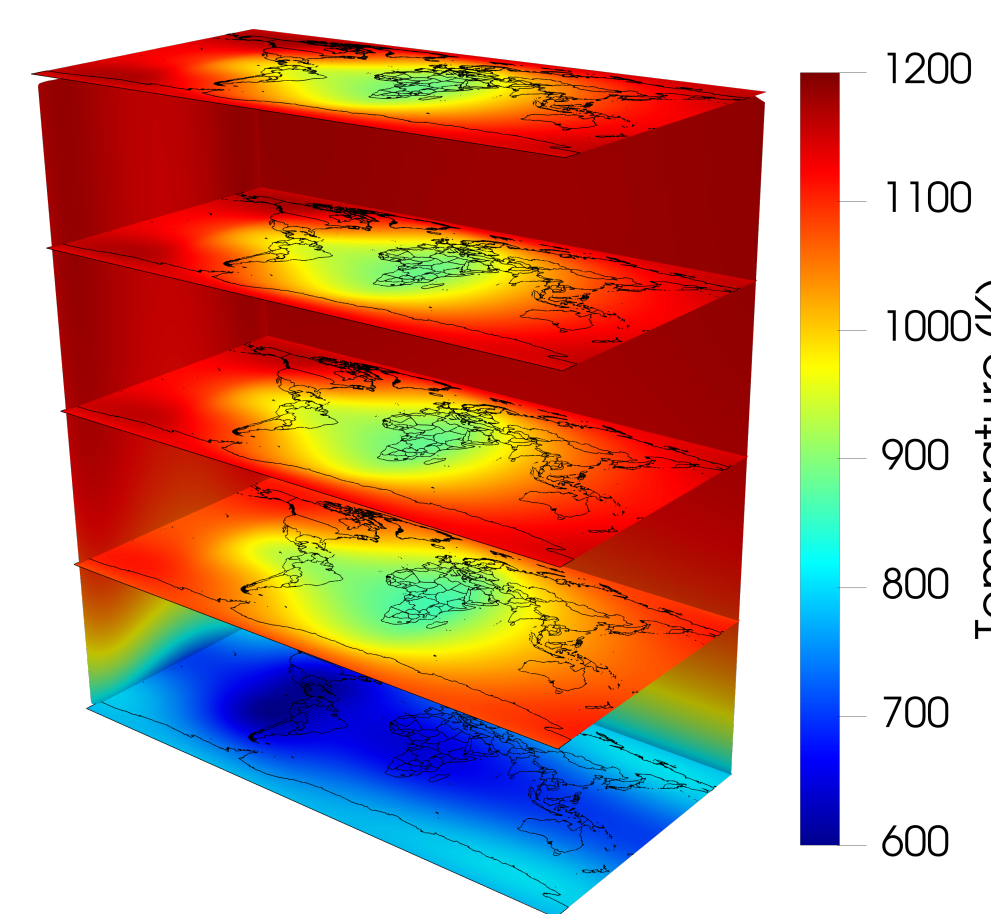
Figure 1: EUV heating source distribution on the globe at a given time.

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4. Computational Study

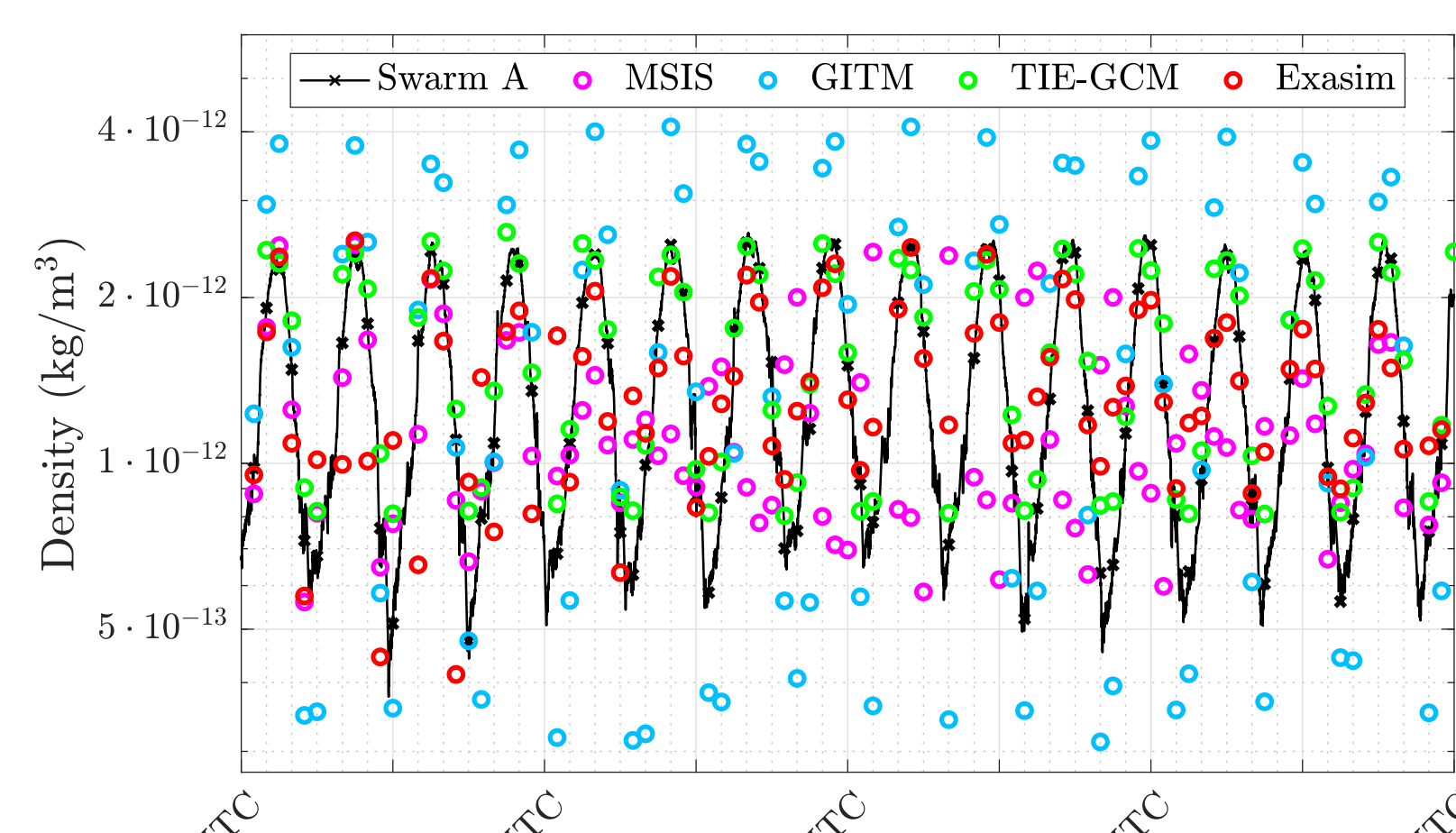
Characterization of thermospheric outputs for 2014 spring equinox.



- **Quadratic approximation** on a cube-sphere mesh with nearly **220k elements** (30M dofs).
- Coherent patterns of density, temperature and velocity.
- Symmetric solution with respect to the Equator line, consistent with solar equinox conditions.

Figure 2: Density estimates at different altitude levels (150, 250, 350, 450 and 550km) at March 22 at midnight (UTC).

Validation with Swarm A satellite data



	MSIS	GITM	TIE-GCM	Exasim
RMS error	0.5325	0.4871	0.1531	0.2528

Figure 3: Density along the Swarm A trajectory (top) and RMS error with respect to accelerometer data [5] (bottom) for Exasim, MSIS [4], GITM [6] and TIE-GCM [7].

- **Error within common ranges** for physics-based thermospheric models.
- Good agreement between **1D and 3D models**.

1D-3D validation

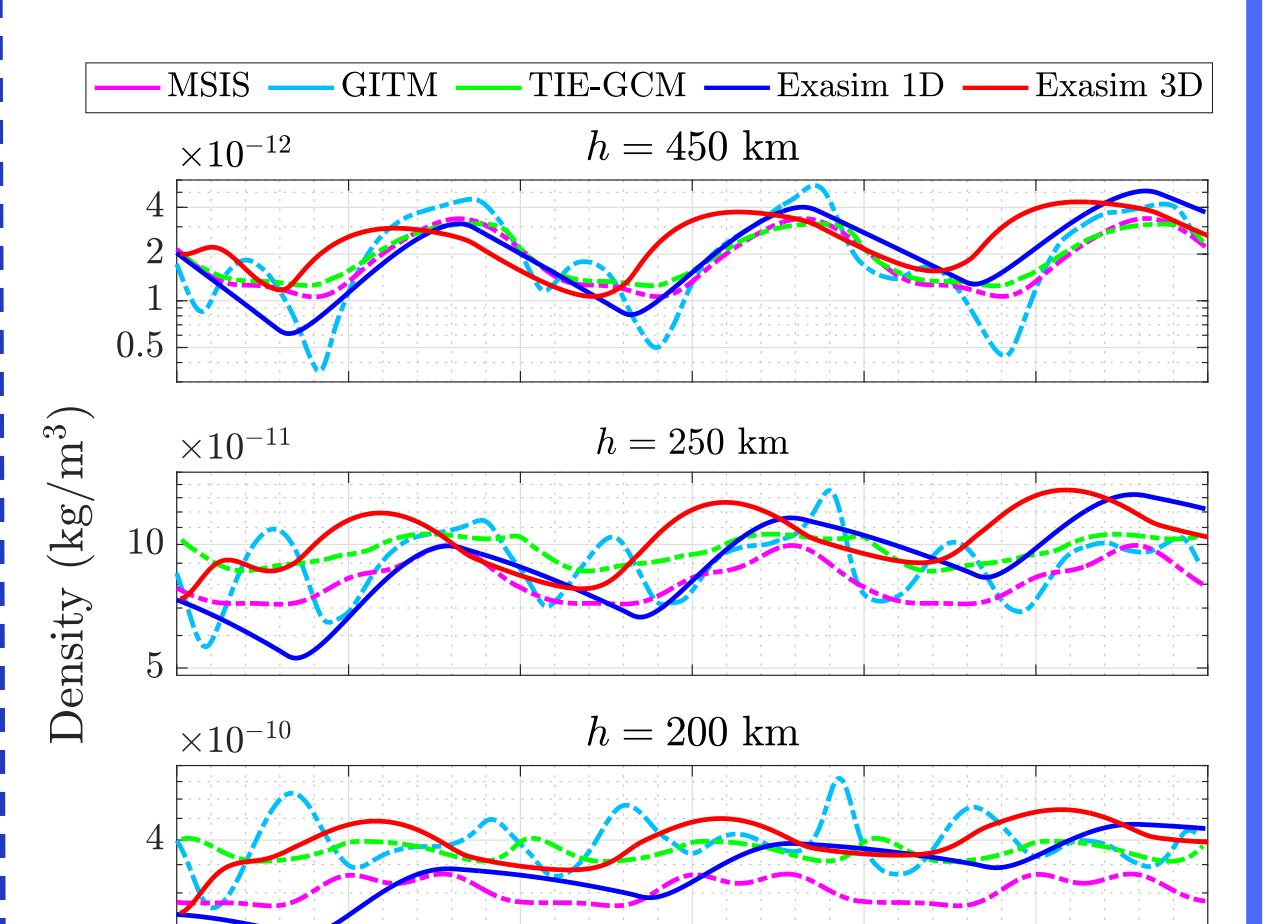


Figure 4: Comparison of the density temporal evolution at different altitudes above Jicamarca, Peru for different thermospheric models during March 2014.

5. Conclusions

High-order DG method for space weather modeling.

- Implicit solver with automatic code generation.
- Matrix-free approach, optimized for GPU computing.

Thermospheric model: neutral gas mixture under EUV heating.

- **Scaling of variables** for multiscale variations and stiffness.
- Close reproduction of **dynamical behavior** of the thermosphere.

Ongoing work

- Ionospheric model development (charged particles, chemistry).
- Integration with uncertainty quantification and data assimilation.

References

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